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07.03.2019 20:45 Coronae (C) are important wavelength-dependent photophysical signatures of coronal holes. They are narrow, steep, and short-lived with considerable width in the coronal hole (ch) loops. The temperature of the loops is determined from the underlying magnetic field. These loops observed in AIA 1600 Å images generally have temperature  $T > 10$  MK or  $T > 4$  MK. The temperature of coronal loops is in the range  $T = (1.7-5.0)$  MK. The loops are hot only at their bases, but their mid-heights can be cool. Most of them carry cool plasma, but there are loops where a hot plasma is also observed. A hotter loop is related to a larger opening. The temperature varies by at most a factor of 2 within each loop. This suggests that the plasma heating process occurs at the footpoint of the loop and that it is consistent with non-potentiality of the loop field. The temperature varies from loop to loop over a range of at least 30 MK. Some of these loops are coherent, some are not, and some have a continuously different temperature distribution. We find a simple relation between the loop half length and the corresponding peak temperature Archived from the original PDF on 09 August 2013. and presented a rough estimation of the heating rate by using an ion cyclotron resonance heating which is widely accepted for solar coronal plasma heating. The released energy from the loop-length dependent temperature is also considered As a result of this study, a new type of solar coronal plasma modeling has been suggested. The study showed that the temperature of the coronal loop peaks at the loop top. The temperature peak increases with loop lengths. However, loops of different lengths converge and form a flux tube with a single temperature distribution. At the footpoints of these converging loops, the local temperature of the loop is formed by a flux tube of a single temperature distribution which is a function of the loop length. The heating rates for loops of different lengths converging from different footpoints are different, but they converge and merge. By this model, the single, cold plasma structures are formed. We also showed that the ion cyclotron resonance heating would not explain the significantly high temperature observed at the top of long loops. Thus, in addition to the estimation of the heating rate from the loop-length dependent loop temperature, the heating process for long loops is to be considered

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